



INAF ISTITUTO NAZIONALE DI ASTROFISICA

YESTERDAY AND TODAY TRACKING AT TNG, TOMORROW SOXS AT THE TWIN NTT

P. SCHIPANI

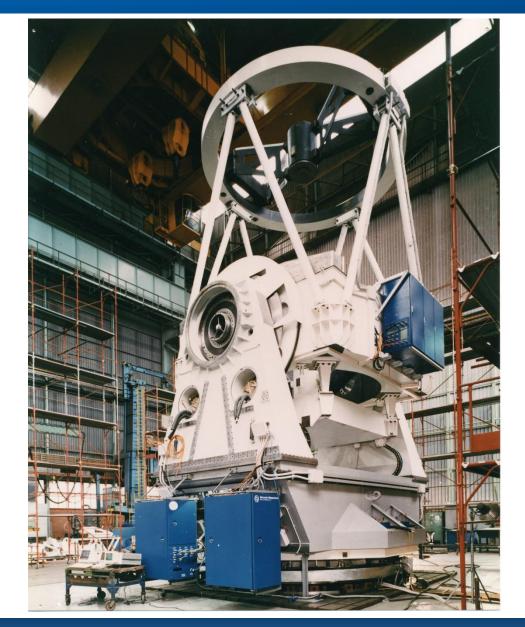
TELESCOPIO NAZIONALE GALILEO: 25 YEARS OF ASTRONOMY IN LA PALMA 19.10.2021



TNG Inaugurations Photo album & Backstage

Ansaldo Milano, December 1994



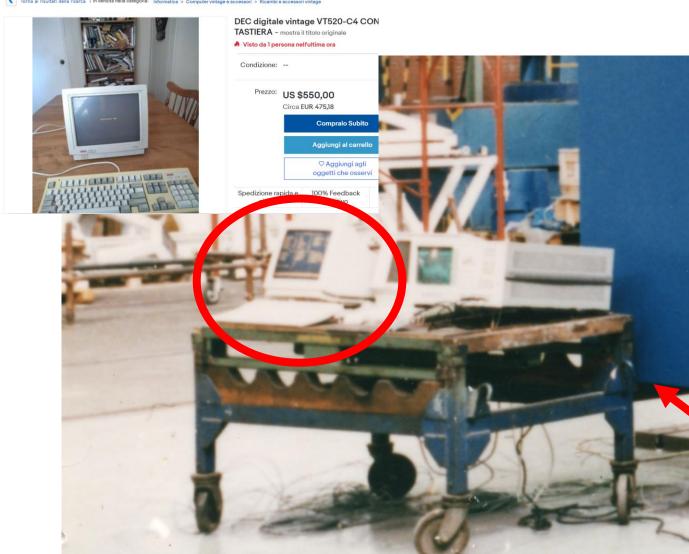


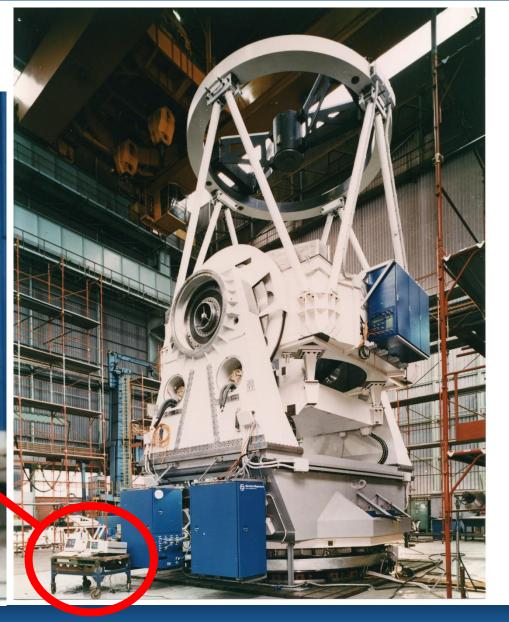
Ansaldo Milano, December 1994



Cerca qualsiasi cosa

Computer vintage e accessori > Ricambi e accessori vintage





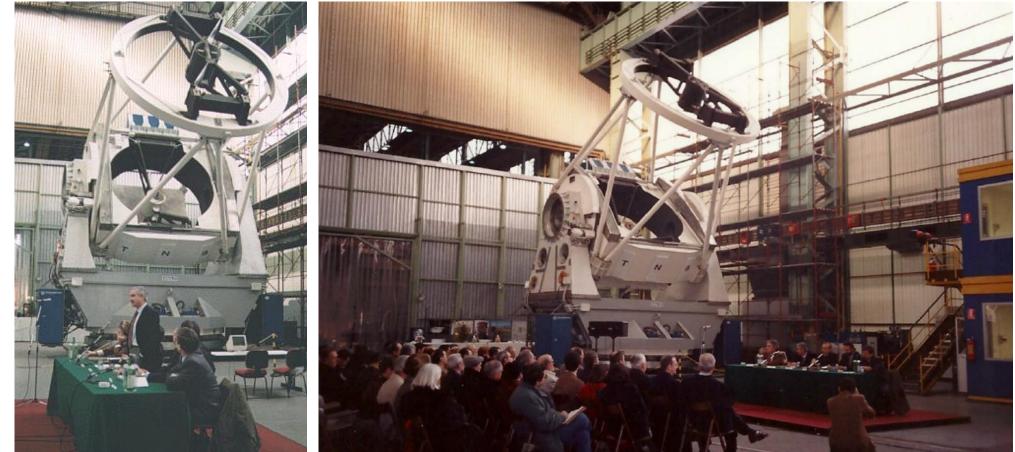
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Ansaldo Milano, December 1994



Show at the inauguration: sinusoidal speed Alt-Az dance...

"The TNG silently and elegantly turned in azimuth and then lowered in elevation as if bowing to the audience in an effortless ballet, and one could not help imagining it inside its dome while pointing some far-off object in the incredibly bright Milky Way shining over the Roque." TNG Newsletter no.9



Position loop closed with encoders in June 1995. First "tracking" of a position trajectory.

La Palma, 1996.06.28 (-1 to Inauguration day)







La Palma, 1996.06.29 TNG Inauguration





La Palma, 1996.06.29 TNG Inauguration



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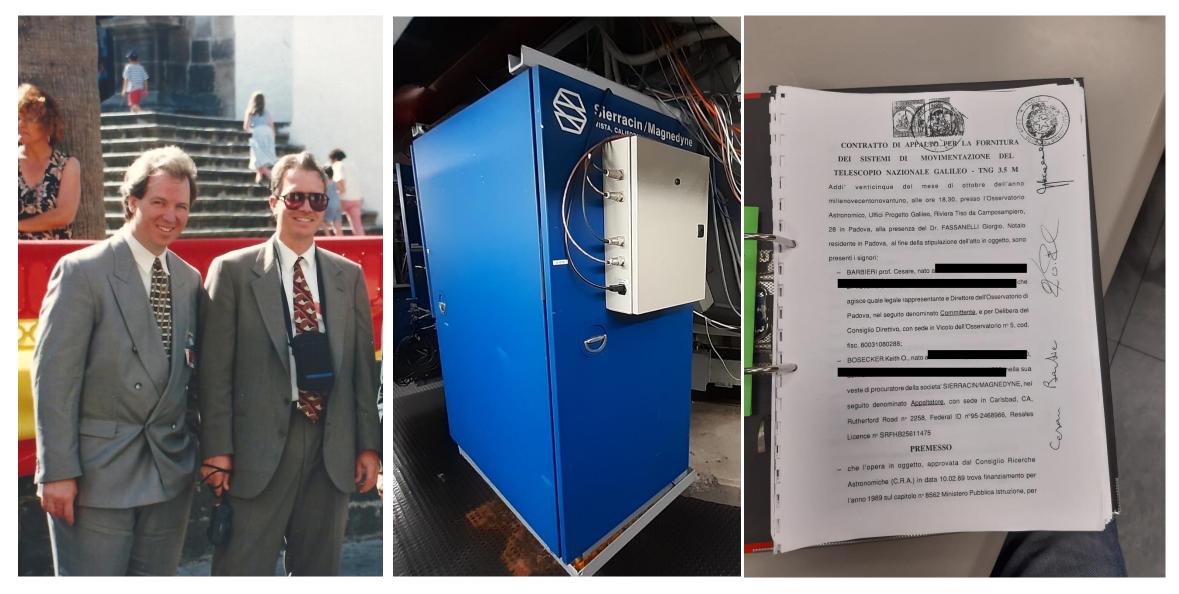
La Palma, 1996.06.29 TNG Inauguration





La Palma, 1996.06.29





My 'Tracking at the TNG'

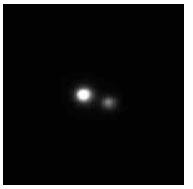


Lot of fun

- □ Heavy work
- □ Unrepeatable week-ends: alone at the TNG

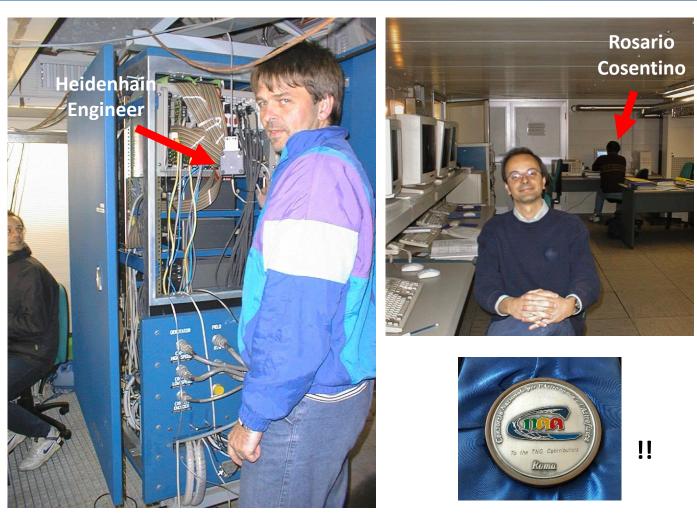
Exciting moments

TNG first light: 9 June 1998 But also the days after...



□ Work completed in 2000

- But keeping always in touch later
- Onsite again in 2003 (good job with Tino)
- □ Fruitful collaboration with Manuel over the years. Now, working together again.
- Axes Control System and SW operational all the nights in 25 years!



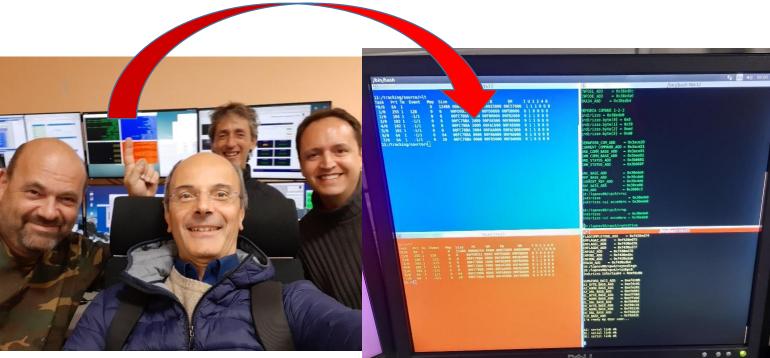
TNG Tracking experience paved the way for the VST Tracking INAF optical telescopes point&track with internal work and know-how

La Palma, 2019-pres.: back to work



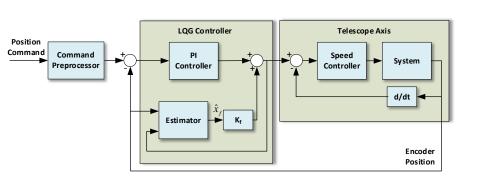
Back to the TNG
TNG Servo Control Redesign
Azimuth, Altitude

N. Gonzalez & Colleagues New electronics Great job ongoing! Congrats



Control loops, hidden but important

New System Identification White Noise **Torque Input** Position Command Position Speed System Controller Controller Encoder d/dt Velocity Output



INAF

Studying new
control strategies
Command shaper
LQG Control

INAF Telescopes Pointing & Tracking is cool





spectrograph from U to H band @ESO-NTT 350-2000 nm *'Similar'* to X-Shooter @VLT
Two arms (VIS + NIR) with partial overlap around 800 nm to crosscalibrate spectra

□Single-object wide band

□R~4,500 (3,500-6,000)

□Acquisition camera to perform photometry ugrizY-V (3.5'x3.5')









Consortium

Institutes from 6 Countries

□ Istituto Nazionale di AstroFisica (INAF), Italy Department of Particle Physics and Astrophysics, Weizmann Institute of Science, Rehovot, Israel Universidad Andres Bello & Instituto Milenio de Astrofisica (MAS), Santiago, Chile FINCA - Finnish Centre for Astronomy with ESO & Turku University, Turku, Finland Queen's University Belfast, UK Tel Aviv University, Israel □ Niels Bohr and Aarhus University, Copenhagen, Denmark

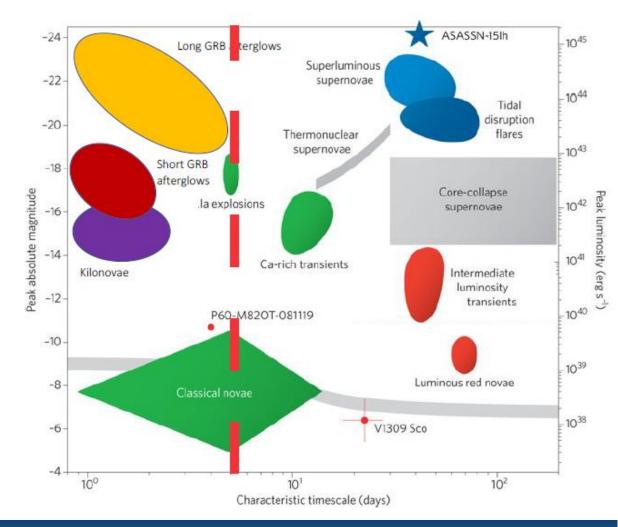






Spectroscopic follow up of transients

- Classification of transients
- □ Supernovae (all flavours)
- Gravitational Wave events
- □ Neutrino events
- Blazars and AGN
- Nuclear transients and Tidal Disruption Events
- GRB and FRB
- Transient X-ray binaries, magnetars, ultraluminous X-ray sources (NS & BH)
- Asteroids and Comets
- □ Young Stellar Objects, stellar variability, exoplanets
- The Unknown





Sinergies

SOXS will have 180 n/yr (for ≥5 yr) ~3,000-4,000 spectra/yr

A spectroscopic machine for the transient sky

□ New deeper survey: Vera Rubin, PanSTARSS, DES, ZTF, ...

- □ Space optical missions: Gaia, EUCLID, ...
- □ Space high-energy missions: Swift, Fermi, SVOM, ...
- □ Radio new facilities: MeerKAT, SKA, ...

UVHE: CTA

□ Messengers: aLIGO-Virgo, KM3Net, ANTARES, ...





NAE



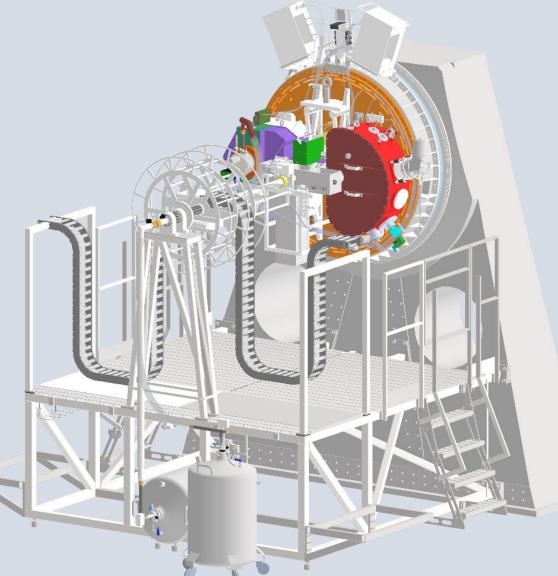


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	UV-VIS	NIR
Spectral range	350-850 nm	800-2000 nm
Resolution (1" slit)	>3600 (≈4500 avg)	5000
Slit widths	0.5 - 1 - 1.5 - 5 arcsec	0.5 - 1 - 1.5 - 5 arcsec
Slit height	12 arcsec	12 arcsec
Detector	e2V CCD44-82 2Kx4K	Teledyne H2RG 2Kx2K
Pixel Size	15 µm	18 µm
Detector Scale	0.28"/pixel	0.25"/pixel

	Camera
Spectral range	360-970 nm
Detector	Andor iKon M-934 1Kx1K
Field of View	3.5'x3.5'
Pixel Size	13 µm
Detector Scale	0.205"/pixel



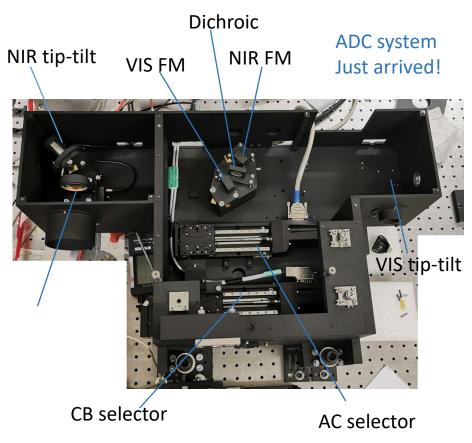






Taking shape

Common Path



Flange

Co-Rotator



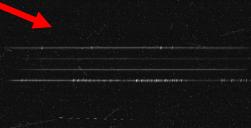
Big problems and delays with COVID

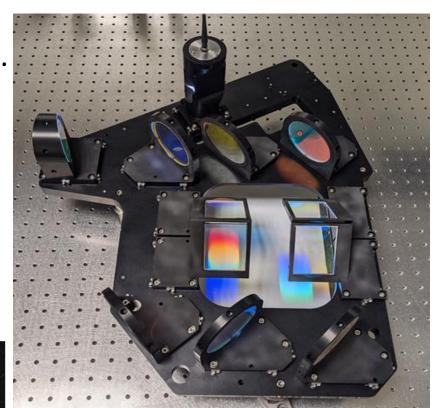


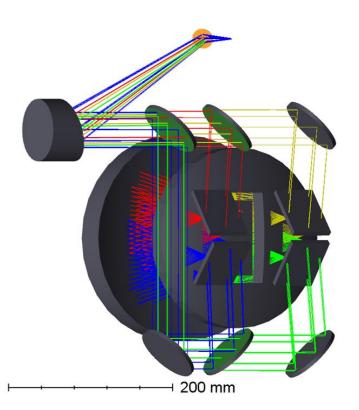


UV-VIS: Multi-Imaging Transient Spectrograph

 Collimated beam is divided to 4 bands using 3 dichroics.
 Each band has its own optimized disperser
 Single camera
 1st order dispersion, *R*~4500 at α_{Lit}.
 4 bands quasi-orders are imaged onto a single 4k×2k CCD.







Quasi-	Wavelength Range	
Order	[nm]	
u	350 – 439.5	
g	427 - 547	
r	527 - 680	
i	664 – 850	





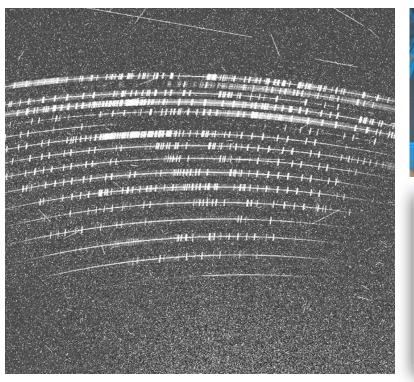
TNG staff responsible of the UV-VIS Detector System





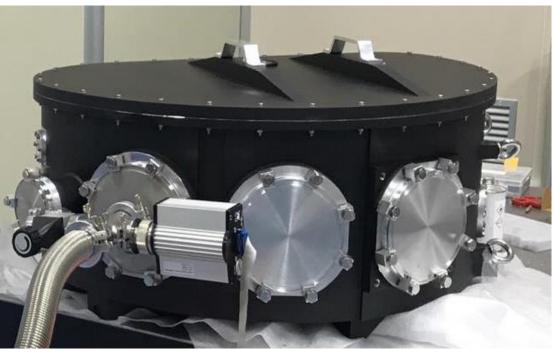


NIR Spectrograph









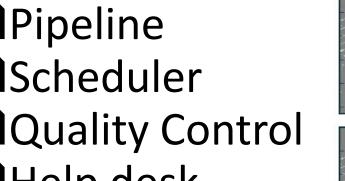


DPipeline **Scheduler Quality Control** Help desk

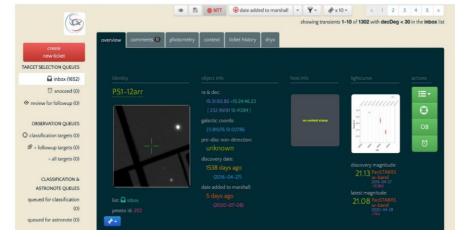
soxs_disp_solution

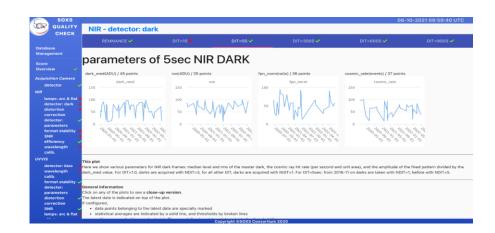
residuals of global dispersion solution fitting - single pinhole mean res: 0.12 pix, res stdev: 0.06 observed arc-line positions (post-clipping)

global dispersion solution



Consortium responsibility for 5 years







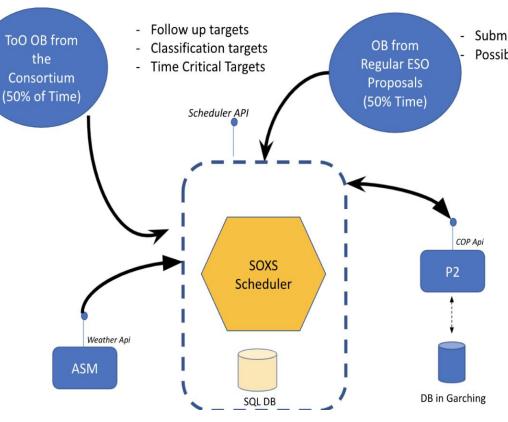


Operations

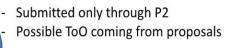


Scheduling

□ Schedule is updated daily Telescope operator on site



P2 system, vOT interface with ESO La Silla 🚞 □Web based app



Marshall Feeders: \Box ZTF, ATLAS, PanStarrs, LSST-Lasair, etc. TNS, Atel, GCN, etc.

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List of OBs

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FUERZA LA PALMA Grazie Fabio